### FacetClumps: A Facet-based Molecular Clump Detection Algorithm

Yu Jiang,  $^{1,2,3}$  Zhiwei Chen,  $^3$  Sheng Zheng,  $^{1,2}$  Zhibo Jiang,  $^3$  Yao Huang,  $^1$  Shuguang Zeng,  $^1$  Xiangyun Zeng,  $^1$  and Xiaoyu Luo $^1$ 

<sup>1</sup> Center for Astronomy and Space Sciences, China Three Gorges University,

8 University Road, 443002 Yichang, China

<sup>2</sup> College of Science, China Three Gorges University,

8 University Road, Yichang, China

<sup>3</sup> Purple Mountain Observatory, Chinese Academy of Sciences,

10 Yuanhua Road, 210023 Nanjing, China

(Received October 30, 2022; Revised May 9, 2023; Accepted May 28, 2023)

Submitted to ApJS

#### ABSTRACT

A comprehensive understanding of molecular clumps is essential for investigating star formation. We present an algorithm for molecular clump detection, called FacetClumps. This algorithm uses a morphological approach to extract signal regions from the original data. The Gaussian Facet model is employed to fit the signal regions, which enhances the resistance to noise and the stability of the algorithm in diverse overlapping areas. The introduction of the extremum determination theorem of multivariate functions offers theoretical guidance for automatically locating clump centers. To guarantee that each clump is continuous, the signal regions are segmented into local regions based on gradient, and then the local regions are clustered into the clump centers based on connectivity and minimum distance to identify the regional information of each clump. Experiments conducted with both simulated and synthetic data demonstrate that FacetClumps exhibits great recall and precision rates, small location error and flux loss, a high consistency between the region of detected clump and that of simulated clump, and is generally stable in various environments. Notably, the recall rate of FacetClumps in the synthetic data, which comprises  $^{13}CO$  (J=1-0) emission line of the MWISP within  $11.7^{\circ} \le l \le 13.4^{\circ}$ ,  $0.22^{\circ} \le b \le 1.05^{\circ}$  and 5 km s<sup>-1</sup>  $\le v \le 35$  km s<sup>-1</sup> and simulated clumps, reaches 90.2%. Additionally, FacetClumps demonstrates satisfactory performance when applied to observational data.

Keywords: radio lines: ISM - ISM: molecules, structure - stars: formation - method: data analysis - techniques: image processing

#### 1. INTRODUCTION

Molecular clouds contain a significant proportion of gas and dust, and are the birthplace of many prominent young objects. Star formation processes take place on the scale of giant molecular clouds or even smaller. (Shu

Corresponding author: Zhiwei Chen

zwchen@pmo.ac.cn

Corresponding author: Sheng Zheng

zsh@ctgu.edu.cn

et al. 1987; Testi & Sargent 1998). Great efforts have been invested in characterizing the feature of molecular gas (e.g. Blitz & Shu 1980; Gammie et al. 2003), deriving the stellar initial mass function (e.g. Chabrier 2003; Alves et al. 2007; Lodieu 2013), and observing the local star formation rate (e.g. Narayanan et al. 2007; Narayanan et al. 2008). Carbon monoxide surveys are a crucial way in unveiling the mysteries of stellar formation, stellar evolution, and galactic structure (e.g. Dame et al. 1987, 2001; Sanders et al. 1986; Lee et al. 2001; Li et al. 2018; Su et al. 2019). Giant molecular clouds often exhibit substructures, such as filaments, clumps,

and cores (Blitz & Williams 1999), while faint sources can be easily obscured by noise. A critical challenge in many research projects is how to accurately segment giant molecular clouds and detect faint targets in the data from carbon monoxide surveys.

Some of the available clump-finding algorithms include GaussClumps (Stutzki & Guesten 1990), ClumpFind (Williams et al. 1994), ReinHold (Berry 2013), FellWalker (Berry 2013), Local Density Clustering (LDC) (Luo et al. 2022), and ConBased(Jiang et al. 2022). The location of peaks is a crucial aspect of GaussClumps, ClumpFind, ReinHold, and LDC. Gauss-Clumps starts from the brightest peak in the data cube to fit the ellipsoid clump, subtracts the fitting clump, and then performs the fitting from the brightest peak in the residuals. ClumpFind contours the data array at many different levels, with a peak being an isolated contour, then works from the highest contour levels to a specified minimum contour level. ReinHold and LDC both rank the data in descending order, with those at the top and above a minimum intensity being considered as potential peaks; For ReinHold, a peak is deemed significant if the pixels spanned by the peak along any one dimension are greater than a specified minimum number; For LDC, a peak is deemed significant if no pixels with greater intensity exist within a specified neighborhood. FellWalker and ConBased do not rely on alternative peaks but instead take into account the relationships between the nearest peaks. FellWalker ascends the line of greatest gradient until a peak is reached, then jumps to the pixel with the highest value in an extended neighborhood to identify clumps, and merges adjacent clumps if their peak-dependent dip is less than a specified value. ConBased divides signals into small regions and merges them from the regions with the smallest volume, using a merging rule based on connectivity, peak distance and intensity differences, and volume.

Molecular clumps are irregular in shape and are characterized by faint gas enveloping a denser central source. To address the two main objectives of molecular clump detection, namely identifying the location and region of molecular clump, we propose a novel algorithm called FacetClumps. The location of peaks, particularly those of faint clumps, is readily impacted by noise. To identify the location of the denser central source and reduce the reliance on the peak in the detection process, FacetClumps utilizes morphology (e.g. Serra 1982; Sinha & Dougherty 1992; Koskinen et al. 1991; Jiang et al. 2022) to extract signal regions from the original data, and incorporates the Gaussian Facet model (Haralick 1984; Qiang & Haralick 2002; Brejl & Sonka 2000; An 2007) and extremum theory of multivariate func-

tion to locate clump centers in the signal regions. A single molecular clump is relatively smooth and continuous; however, FellWalker may detect multiple distinct components as a single clump, while LDC obtains the connectivity by selecting the subpart with the largest volume from the potential clumps and discarding the smaller, discontinuous subparts. To improve the accuracy of regional segmentation, FacetClumps utilizes a gradient-based method (Berry 2013) to segment the signal regions into local regions, and then applies a connectivity-based minimum distance clustering method to cluster the local regions to the clump centers. To improve the adaptability, the parameters of FacetClumps are automatically adjusted according to different local situations, and are optimized for detecting faint and overlapping clump.

We illustrate the processes and details of FacetClumps by combining text and schematic diagrams in Section 2. In Section 3, we determine suitable values for FacetClumps parameters in simulated clumps, and compare its performance with that of other algorithms in simulated and synthetic clumps. We then apply FacetClumps to observational data, and present the results in Section 3. In Appendix A, we conduct experiments on larger synthetic data with different signal densities. In Appendix B, we analyze the performance of FacetClumps in different resampled synthetic data. Finally, we summarize our work in Section 4.

#### 2. THE FACETCLUMPS ALGORITHM

FacetClumps primarily consists of four sub-processes. The first sub-process is signal region extraction based on morphology, which includes threshold segmentation, opening, dilation, and connected domain labeling of the original data, followed by combining the connected domain with the original data to acquire the signal regions. The second sub-process is clump center detection based on the Facet model. The Gaussian Facet model operators are convolved with the signal data, resulting in a fitting surface and corresponding fitting coefficients. The first and second derivatives of the fitting surface are then calculated from these coefficients. The extremum determination theorem is employed to identify potential maxima by utilizing the first derivatives and the eigenvalues of the Hessian matrix constructed from the second derivatives. To account for the interference of noise, the maximum regions which may contain maxima are extracted by adaptively adjusting the thresholds of the first derivatives and the eigenvalues. The centroid of a maximum region is taken as the clump center. The third sub-process is local region segmentation based on local gradients. In the smallest neighborhood of voxeles, each

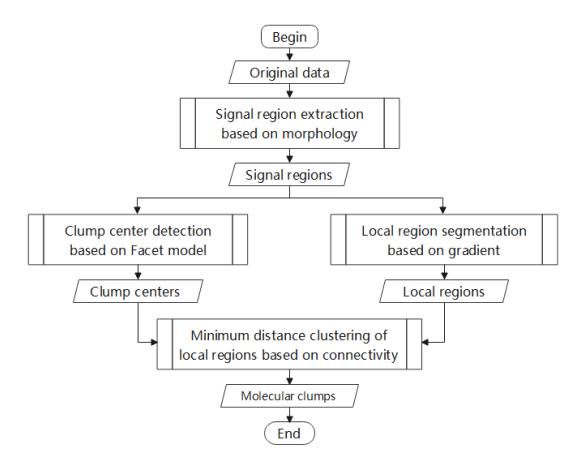


Figure 1. Flow chart. FacetClumps consists of four sub-processes: (1) signal region extraction based on morphology, (2) clump center detection based on the Facet model, (3) local region segmentation based on the gradient, and (4) minimum distance clustering of local regions based on connectivity.

signal region is segmented into local regions by ascending along the highest gradient of intensity (Jiang et al. 2022). The fourth sub-process is minimum distance clustering of local regions based on connectivity. Matching clump centers with local regions, each matched region is regarded as the target region of a clump. The connectednearest local regions of a target region are then merged into the target region, forming clumps. This merging process continues until all local regions are clustered.

Finally, the statistics of each clump in pixel and WCS coordinate systems are collected in two tables, and the regional information of each clump is recorded in a mask. The flow of FacetClumps is shown in Figure 1. FacetClumps can be applied to two-dimensional (position-position space, hereafter PP) and three-dimensional (position-position-velocity space, hereafter PPV) observational data, but we will mainly introduce the PPV flow. We have shared the code on

Github under a permissive MIT license<sup>1</sup>, made it publicly available as a Python package called FacetClumps<sup>2</sup>, and deposited the latest version to Zenodo (Jiang 2023). We warmly welcome community contributions to its optimization.

## 2.1. Signal region extraction based on morphology

To enhance the efficiency and robustness of FacetClumps, the original data is preprocessed based on morphology. As illustrated in Figure 2(a), there are five clumps, three of which are overlapping. Figure 2(b) shows the PPV view of Figure 2(a), and Figure 2(c) displays the clumps after introducing noise.

Firstly, a free parameter Threshold (typically  $2 \times RMS$ , where RMS is the noise RMS of the data) is selected to binarize the original data. Voxel values greater

 $<sup>^{1}\;</sup> https://github.com/JiangYuTS/FacetClumps$ 

<sup>&</sup>lt;sup>2</sup> https://pypi.org/project/FacetClumps/

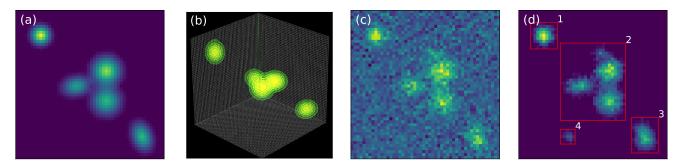


Figure 2. (a) The original noise-free data; (b) The original noise-free data with a PPV view; (c) Noisy data; (d) The signal regions, with a red box being the smallest enclosing rectangle of a region.

than the threshold are labeled as binarization mask. Secondly, a morphological opening operation (Chen & Haralick 1995) is performed on the binarization mask to obtain an opening mask. The opening operation separates the boundaries of the regions from the noise, which eliminates noise points and has a negligible impact on the size of the regions. Thirdly, to reduce the loss of flux and enhance the detection of faint clumps, a morphological dilation is performed on the opening mask, and the signals below the noise level and above the threshold of the dilated voxels are removed to obtain a dilation mask. The kernels of the opening and dilation operators are ball operators of radius one. Fourthly, a connected domain operator is applied to the dilation mask to obtain a connected domain mask.

Connectivity is a relationship between neighborhoods. In PPV, there are three distinct types of connectivity (Williams et al. 1994): TypeI, TypeII, and TypeIII. Each type requires that the maximum number of orthogonal hops of a voxel considered to be a neighbor are 1, 2, and 3, respectively (Wu et al. 2005; Wu et al. 2008). A schematic diagram of this is shown in Figure 1 of Williams et al. (1994).

The connected domain masks with intensity are indicative of signal regions, where the connectivity type is TypeIII, as depicted in Figure 2(d). It is evident that the signal region in box 2 consists of three overlapping clumps, and the signal region in box 4 is composed of noise.

# 2.2. Clump center detection based on Facet model

# $2.2.1. \ \textit{Basic principle of PPV Gaussian Facet model}$

Facet model utilizes the concept of sub-sections to obtain the most accurate analytic function in a regular region centered around a specific pixel, using the least-squares fitting method. Wu & Wee (1999) extended the two-dimensional directional derivative edge detector to three-dimensional. Boomgaard & Weijer (2003) investigated the Gaussian weighted Facet model from

the linear scale-space category. Unlike the traditional Haralick Facet model, the Gaussian Facet model employs the weighted least-squares fitting method to obtain the operators of the model. The Gaussian function serves as the weight, indicating the varying importance of different sampling points in determining the final operators. The weight is higher when a sampling point is closer to the center of the region. An (2007) derived the three-dimensional edge detection operator based on the Gaussian Facet model and applied it to three-dimensional subvoxel surface detection, achieving good performances.

For simple images, the intensity function can be approximated by a piecewise constant or piecewise bivariate linear function. For complex images, higher-order polynomials should be chosen. In this paper, we use a polynomial of ternary cubic with integer coefficients to establish the PPV Gaussian Facet model. The original molecular clump surface I(x, y, z) can be approximated by a linear combination of a set of bases  $g_i(i = 1, 2, ..., 20)$ , and the approximation function f(x, y, z) is defined as:

$$f(x, y, z) = \Phi \mathbf{a}$$

$$\mathbf{a} = (a_1, a_2, \dots, a_{20})^{\mathrm{T}}$$

$$\Phi = (g_1, g_2, \dots, g_{20}) = (1, x, y, z, x^2, y^2, z^2, xy, xz, yz, xyz, xyz, xz^2, x^2y, yz^2, x^2z, y^2z, x^3, y^3, z^3)$$

(1)

The least squares fitting minimizes the difference  $\varepsilon$  of the molecular clump surface I(x, y, z) and approximation function f(x, y, z):

$$\varepsilon = \iiint_{\Omega} (I(x, y, z) - f(x, y, z))^2 W(x, y, z) dx dy dz$$
(2)

where W(x, y, z) is the window function defining the locality of the model fitting. The Gaussian Facet model employs the Gaussian window function:

$$W(x, y, z) = \frac{1}{2\pi s^2} exp(-\frac{x^2 + y^2 + z^2}{2s^2})$$
 (3)

where s represents the window radius. The window scale SWindow is a free parameter, and  $s = \lfloor SWindow/2 \rfloor$ . The optimal parameter vector  $\boldsymbol{a}$  is obtained by projecting the function f onto the subspace spanned by the basis functions in  $\Phi$ . The inner product in this function space is expressed as:

$$\begin{aligned} p*q &\equiv \langle p,q \rangle W \\ &= \iiint_{\Omega} p(x,y,z)q(x,y,z)W(x,y,z)dxdydz \end{aligned} \tag{4}$$

where p, q are arbitrary functions of three variables. So the difference  $\varepsilon$  in (2) is rewrited as (5):

$$\varepsilon = \langle I - \Phi \boldsymbol{a}, I - \Phi \boldsymbol{a} \rangle W$$

$$= (\langle I, I \rangle - 2\langle I, \Phi \boldsymbol{a} \rangle + \langle \Phi \boldsymbol{a}, \Phi \boldsymbol{a} \rangle) W$$

$$= I^{T} * I - 2\boldsymbol{a}^{T} \Phi^{T} * I + \boldsymbol{a}^{T} \Phi^{T} * \Phi \boldsymbol{a}$$
(5)

Taking the derivative of (5) for  $\boldsymbol{a}$ , setting  $\frac{\partial \varepsilon}{\partial \boldsymbol{a}} = 0$ , and solving for  $\boldsymbol{a}$ , we obtain:

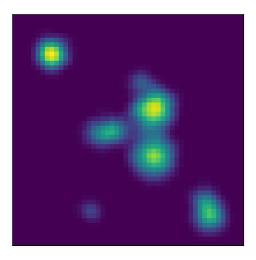
$$\mathbf{a} = (\Phi^{\mathrm{T}}\Phi)^{-1}\Phi^{\mathrm{T}} * I = K * I$$
 (6)

The convolution operators K can be obtained by substituting  $\Phi, W$  into (6) and utilizing the weighted inner product of (4). The convolution of K with the original surface yields the coefficient  $\boldsymbol{a}$  of its fitting surface. The continuous form of the inner product not only provides higher accuracy than the discrete form, but also incorporates the window scale into the definitive  $\boldsymbol{a}$ , which is advantageous for multi-scale analysis.

# 2.2.2. Combine multivariate function maximum determination theorem and Facet model

Let a multivariate real function f(x, y, z) have the continuous second derivative in the neighborhood of point  $P_0(x_0, y_0, z_0)$ . If  $\frac{\partial f}{\partial \mathbf{x}}|_{(x_0, y_0, z_0)} = 0$ , and H constructed by the second derivative of f is a negative definite matrix,  $f(x_0, y_0, z_0)$  is a maximum. The necessary and sufficient condition for a matrix to be negative definite is that all its eigenvalues are less than zero.

Convolve the data cube of the signal regions in Section 2.1 with the operators K in Section 2.2.1 to derive the first and second derivatives of the fitting surface. The fitting surface is shown in Figure 3. The first derivatives  $(F_x, F_y, F_z)$  of  $f(x, y, z)_{(x,y,z)=(0,0,0)}$  are obtained by (7):



**Figure 3.** The integral diagram of the fitting surface of the singal regions shown in Figure 2(d).

$$F_x = \frac{\partial f}{\partial x} = a_2, F_y = \frac{\partial f}{\partial y} = a_3, F_z = \frac{\partial f}{\partial z} = a_4$$
 (7)

Formula (8) reveals the second derivatives of the central point of neighborhood, which are used to construct the Hessian matrix, as demonstrated in (9).

$$\frac{\partial^2 f}{\partial x^2} = 2a_5, \frac{\partial^2 f}{\partial y^2} = 2a_6, \frac{\partial^2 f}{\partial z^2} = 2a_7,$$

$$\frac{\partial^2 f}{\partial xy} = a_8, \frac{\partial^2 f}{\partial xz} = a_9, \frac{\partial^2 f}{\partial yz} = a_{10}$$
(8)

$$H = \begin{pmatrix} 2a_5 & a_8 & a_9 \\ a_8 & 2a_6 & a_{10} \\ a_9 & a_{10} & 2a_7 \end{pmatrix}$$
 (9)

According to the multivariate function maximum determination theorem and considering the interference of noise, any one of the first derivatives around the center of a clump is near zero, and the eigenvalues of the Hessian matrix around the center of a clump are less than zero. As a result, during clump center detection, we do not simply search for maximum values of the fitting surface directly, but instead extract a certain range of maximum regions.

#### 2.2.3. Locate the clump centers

We define operations OP1, OP2, OP3, and OP4 to obtain adaptive thresholds for the first derivatives and eigenvalues, and use them to search for the maximum

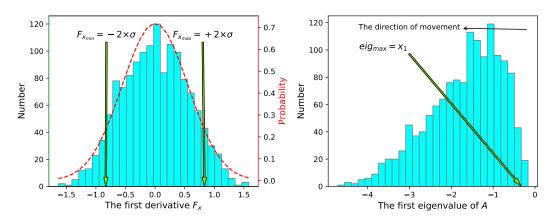
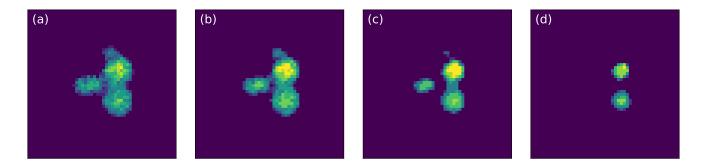


Figure 4. Distribution of the first derivative  $F_x$  and the first eigenvalue of the region numbered 2 in Figure 2(d). The distribution of  $F_x$  is fitted with the Gaussian function. The thresholds of the first derivative  $F_x$  are the double negative and positive standard deviations  $(-2 \times \sigma \text{ and } +2 \times \sigma)$  of  $F_x$ . Coordinates with  $F_x$  value between the thresholds are used to judge the maximum regions. Similarly, the thresholds of  $F_y$  and  $F_z$  are determined in the same manner as  $F_x$ . The initial threshold of the first eigenvalue is the x-coordinate of the first bin on the right, and the x-coordinate is shifted one bin to the left after each recursion. Eigenvalues that are less than the threshold are taken into consideration for the judgment. Similarly, the thresholds of the second and third eigenvalues are determined in the same manner as the first eigenvalue.



**Figure 5.** Some of the recursion masks of the region numbered 2 in Figure 2(d). (a) The sub-maximum region after the first recursion. (b) The sub-maximum region after the second recursion. (c) The sub-maximum regions separating one of the clumps and the maximum region of the noise cluster. (d) The sub-maximum regions separating the last two clumps.

regions in each signal region. Figure 4 depicts the distribution of the first derivatives  $F_x$  and the first eigenvalues of the signal region which is composed of three overlapping clumps as shown in box 2 of Figure 2(d).

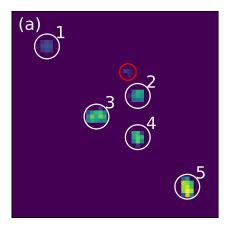
Operation OP1: Determine the preselected maximum regions using the first derivatives. A Gaussian function is used to fit the distribution of the first derivative, and its mean is around zero. The standard deviation of the first derivative values is computed for each region, and the thresholds are negative two standard deviations and positive two standard deviations. Voxels whose values fall within these thresholds are selected as candidates for the preselected maximum region.

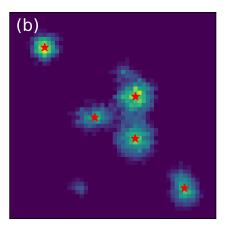
Operation OP2: Determine the preselected maximum regions using the eigenvalue. To better accommodate the distribution characteristics of different regions, the eigenvalues are binned according to the number of bins (Bins), which is calculated by (10). The threshold is the x-coordinate of the ith bin on the right side of the distribution, as shown in Figure 4. (The initial value of i is minus one, and after each recursion, which will be described in operation OP4, the value decreases by one.) Voxels whose values are less than the threshold are selected as the candidates for the preselected maximum region.

$$Bins_i = |KBins \times \ln Vsr_i|$$
 (10)

where, KBins is a free parameter,  $Vsr_i$  is the volume of the *i*th signal region, i = 1, 2, ..., n1, with n1 being the number of signal regions.

Operation OP3: Determine the connected maximum regions using the preselected maximum regions. Both the criteria for the first derivatives and eigenvalues must be simultaneously satisfied. Voxels in the preselected





**Figure 6.** (a) The integrated intensity of the final maximum regions; the white circles represent the valid maximum regions. (b) The clump centers detected by FacetClumps; the background is the signal regions and the red asterisks denote the locations of the clump centers.

sets that meet the criteria are marked as one. The connected domain operator is then applied to the marked data to identify connected domains, and each connected domain is considered to be a maximum region. To improve the ability to detect faint and overlapping clumps, the connectivity type used is TypeI.

Operation OP4: Determine sub-maximum regions using recursion. Figure 5(a) illustrates that a maximum region extracted from the crowded signal region may still contain multiple clumps. To separate these overlapping clumps, recursion operations OP1, OP2, and OP3 are applied to each maximum region if its area in the spatial direction exceeds the parameter SRecursionLB or its length in the velocity channels exceeds the parameter SRecursion V to extract sub-maximum regions. During the recursion process, the Bins of each sub-region is the same as its signal region, the thresholds of the first derivatives and eigenvalues for each subregion are updated as described in operations OP1 and OP2, and the connected sub-maximum regions are determined as described in operation OP3. It can be seen from Figure 5, the overlapping clumps are successfully separated after recursions.

The resulting sub-maximum regions that have undergone a recursion are displayed in Figure 6(a). Some small sub-regions are commonly caused by noise, as indicated by the red circle in Figure 6(a). To determine a valid maximum region, the formula in (11) can be used.

$$Region_j = \begin{cases} True, & \text{if } Vmr_j \ge \log Vsr_i \\ False, & \text{else} \end{cases}$$
 (11)

where,  $Region_j$  represents the jth maximum sub-region,  $Vmr_j$  is the volume of  $Region_j$ , j = 1, 2, ..., n2, with

n2 being the number of maximum sub-regions,  $Vsr_i$  is the volume of the *i*th signal region, i = 1, 2, ..., n1, with n1 being the number of signal regions.

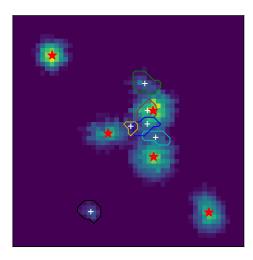
After performing the operations mentioned above, all valid maximum regions are obtained. These regions are combined with the fitting surface to extract intensity information. The centroid of each valid maximum region is calculated using equation (12). This centroid represents the location of a denser central source and is recorded as a clump center. The resulting clump centers are shown in Figure 6(b) and are found to be consistent with the simulated central coordinates.

$$Cen = \frac{\sum_{k=1}^{n3} f[\boldsymbol{u_k}] \cdot \boldsymbol{u_k}}{\sum_{k=1}^{n3} f[\boldsymbol{u_k}]}$$
(12)

where, Cen represents the centroid of a valid maximum region, f is the fitting surface,  $u_k$  is the coordinate of k,  $u_k = \{x, y, z\} = U[k]$ , and U is the coordinate set of a valid maximum region, n3 is the number of the coordinates.

#### 2.3. Local region segmentation based on gradient

The approach of segmenting the signal region based on local gradients has been widely applied and has been shown to be effective (Berry 2013; Jiang et al. 2022). FacetClumps utilizes a similar approach to identify local regions and then merges them together. It differs from the FellWalker in that, when the path reaches a local maximum, FacetClumps does not continue to search for voxels with higher intensity in a larger neighborhood. It differs from the ConBased in that FacetClumps uses the centroid of a local region as the reference position for subsequent clustering.



**Figure 7.** The PP view of some of the local regions and local centers. Each outline delineates the boundary of a local region, the white plus signs denote the local centers, and the red asterisks denote the clump centers.

Starting at any signal region, a random voxel is selected as the center of a box whose size is  $3\times3\times3$  voxels, then calculate the gradients of intensity between it and each neighbor in the box. Move to the neighbor with the highest gradient, which is regarded as the center for the next movement. The way of finding the next central location is described as (13). The moving progress is repeated until the highest gradient is less than zero.

$$u_n = \begin{cases} u_0, & \text{if } max(I[u_i] - I[u_0]) < 0 \\ U[i_{max(I[u_i] - I[u_0])}], & \text{else} \end{cases}$$
(13)

where,  $u_n$  represents the next central coordinate. max() stands for the maximal value of its arguments. I is intensity map, U is the set of neighbor coordinates in a box.  $u_0$  is the central location of U,  $u_i \in U$ , i = 1, 2, ..., 26.  $I[u_i]$  is the intensity of  $u_i$ , and  $I[u_0]$  is the intensity of  $u_0$ .  $U[i_{max(I[u_i]-I[u_0])}]$  is the neighbor coordinate corresponding to the maximum intensity.

All voxels that have been traversed the movement process are recorded as a path, and the end of each path is recorded as a local maximum. Once a local maximum is reached, new local maxima are sought from the unvisited voxels until all voxels in all signal regions have been searched. The voxels on the paths leading to the same local maximum form a local region, and the connectivity type of each local region meets TypeIII. The centroid of each local region is calculated and recorded as a local center. By applying this approach, it is possible to obtain all local centers and regions for each signal

region. Examples of partial local centers and local regions for signal regions 2 and 4 in Figure 2(d) are shown in Figure 7.

# 2.4. Minimum distance clustering of local regions based on connectivity

The following process relies on the signal regions, clump centers, local centers, and local regions, the purpose of which, is to let local regions be clustered to their corresponding clump centers with precision. A local region is deemed to match a clump center if the coordinates of the clump center fall within the region. In PPV space, the distance between two voxels is measured by (14).

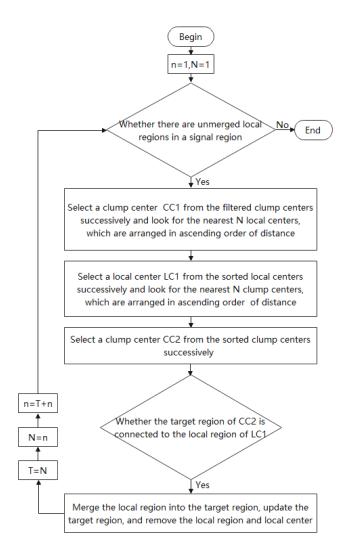
$$Dist = \sqrt{\frac{\Delta x^2 + \Delta y^2}{FwhmBeam^2} + \frac{\Delta z^2}{VeloRes^2}}$$
 (14)

where, Dist represents the distance between  $(x_1, y_1, z_1)$  and  $(x_2, y_2, z_2)$ , FwhmBeam is the full width at half maximum (FWHM) of the instrument beam, and VeloRes is the velocity resolution of the instrument.

Firstly, the clump centers match with signal regions to filter out local regions that cannot be clustered to any one clump center. If the match fails, the signal region and its local regions and local centers will be removed. This commonly occurs when the signal region is composed of noise, as shown by the black contour in Figure 7.

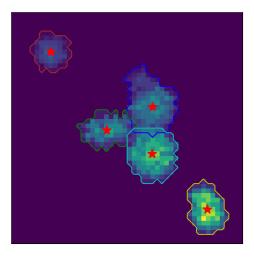
Secondly, clump centers match with local regions to obtain filtered clump centers and target regions. a match is successful, the matched local region is rerecorded as the target region of the matched clump center. The matched local region and its local center will be removed from the record obtained in Section 2.3. If a match fails, this means that a local region may correspond to more than one clump center, due to the difference in connectivity types between the maximum region shown in Figure 6(a) and the local region shown in Figure 7. The local region matched by the failed clump center has already been matched with another clump center and become a target region. In this scenario, to choose a more appropriate clump center for the target region, one of the failed clump center and the matched clump center that is closer to the local center of the target region will be retained, and the other one will be removed. After all the matches have been made, the remaining clump centers are recorded as filtered clump centers, each of which is associated with a target region.

Thirdly, the filtered local regions are merged to the target regions. Connectivity is the necessary condition for merging local regions. In the connected case, a filtered local region is merged into a target region whose clump center is closest to its local center. Since different



**Figure 8.** Diagram of minimum distance clustering of local regions based on connectivity. N is a continuously updated member of the Fibonacci sequence, and n and T are temporary variables. Connectivity is the necessary condition for merging. In the case of connectivity, a filtered local region is merged into a target region whose clump center is closest to its local center.

signal regions are disconnected, the following operations are performed in each signal region. When there are unmerged local regions in any signal region, we introduce a three-layer main loop to merge these local regions, as shown in Figure 8. In the first layer, a clump center of filtered clump centers, denoted as CC1, is selected in the order obtained by recursion to search for the N nearest local centers, which are arranged in ascending order of distance. In the second layer, a local center of the sorted local centers, denoted as LC1, is successively selected to search for the N nearest clump centers, which are also



**Figure 9.** The PP view of the integral mask of the detected clumps. Each outline delineates the boundary of a clump, and the red asterisks denote the clump centers.

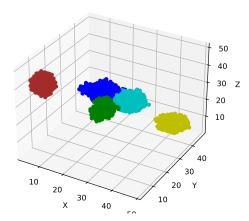


Figure 10. The PPV view of the detected clumps. Different coloured voxels denote different clumps.

arranged in ascending order of distance. In the third layer, a clump center of the sorted clump centers, denoted as CC2, is successively selected to judge whether the local region of LC1 is connected to the target region of CC2, with the connectivity type being TypeIII. If the local region is connected, it will be merged into the target region of CC2 and the target region will be updated. The current local center LC1 and its local region will be removed from the record. If not, the local center LC1 and its local region are retained and await the next judgment. (N is a continuously updated member of the Fibonacci sequence and its initial value is one, which corresponds to the beginning of the sequence. Following each three-layer iteration, the position index of N increases by one. The updated values of N can also be

**Table 1.** The input parameters of FacetClumps.

Parameters	Explanation				
RMS	The noise RMS of the data.				
Threshold	The minimum intensity used to truncate the signals.				
SWindow	The scale of the window function, in pixels.				
KBins	The coefficient used to calculate the number of eigenvalue bins, see formula (10).				
FwhmBeam	The FWHM of the instrument beam, in pixels.				
VeloRes	The velocity resolution of the instrument, in channels.				
SRecursionLBV	The minimum area of a region in the spatial direction (SRecursionLB) and the minimum length of a region in the velocity channels (SRecursionV) when a recursion terminates. The region of a clump also need to satisfy the conditions. See formula (B1), in pixels.				

**Note.** SRecursionLBV consists of SRecursionLB and SRecursionV, i.e. [SRecursionLB, SRecursionV]. The relationship between SRecursionLBV and FwhmBeam and VeloRes is presented in Appendix B.

inferred from Figure 8. If there are fewer items than N, all of them will be searched.) This approach can ensure the stability and accuracy of merging while also improving the efficiency of FacetClumps.

The target regions that ultimately satisfy the condition for the parameter SRecursionLBV will be identified. At this point, we obtain all the clump centers and corresponding regions. An integral mask of the result is shown in Figure 9, where each outline delineates the boundary of a clump. PPV graph is shown in Figure 10, in which different coloured voxels denote different clumps. It is evident that the separated clumps are properly delineated and the connected clumps are accurately segmented.

#### 2.5. The parameters

All the output parameters necessary for the clump tables are computed from the input data, the clump centers, and regions. The regional information is identified by a mask, and the index (starting from one) of each clump corresponds to the same number in the mask. The input parameters of FacetClumps are listed in Table 1, while the output parameters of a single clump are presented in Table 2.

The size (Berry 2013; Jiang et al. 2022) of a clump is defined as (15):

**Table 2.** The output parameters of a clump.

Parameters	Explanation				
PeakI	The peak intensity.				
PeakL	A vector of the peak location, in pixels.				
Cen	A vector of the clump center, see formula (12), in pixels.				
Size	A vector of the sizes, see formula (15), in pixels.				
Sum	The sum of the voxels intensity within the clump.				
Volume	The total number of voxels within the clump, in pixels.				
Angle	The angle of the clump, see formula (16), in degree.				
Edge	Whether the clump touches the edges, 1-Yes, 0-No.				

**Note.** The output tables include tables in both the pixel coordinate system and the WCS coordinate system. The units in the table under WCS coordinate system are consistent with those in the header file.

$$size = \sqrt{\frac{\sum I_i \cdot u_i^2}{\sum I_i} - \left(\frac{\sum I_i \cdot u_i}{\sum I_i}\right)^2}$$
 (15)

where  $I_i$  is the intensity of voxel i minus the minimum intensity in the clump, and  $u_i$  is the coordinate of voxel i

The angle and axis ratio of a clump are calculated by diagonalizing the moment of inertia matrix (Koda et al. 2006), as shown in equation (16):

$$R_{-\theta} \begin{pmatrix} \sum T_i \alpha_i^2 & -\sum T_i \alpha_i \beta_i \\ -\sum T_i \alpha_i \beta_i & \sum T_i \beta_i^2 \end{pmatrix} R_{\theta} = \begin{pmatrix} S_{xx} & 0 \\ 0 & S_{yy} \end{pmatrix}$$
(16)

where T is the velocity-integrated intensity map of a clump,  $T_i$  is the intensity at coordinate  $u_i$ ,  $u_i = \{x, y\}$ ,  $\alpha_i$  and  $\beta_i$  are the Euclidean distances from the clump center to the coordinate  $u_i$  in the l- and b- direction, respectively.  $R_{\theta}$  is a rotation matrix with rotation angle  $\theta$ , which is the angle between the major axis and the negative direction of l, ranging from  $-90^{\circ}$  to  $90^{\circ}$ , and  $\theta$  along the positive direction of b is  $90^{\circ}$ . The axis ratio of a clump is given by the square root of the ratio of the lengths of its major and minor axes, i.e.  $(S_{xx}/S_{yy})^{1/2}$  or  $(S_{yy}/S_{xx})^{1/2}$ .

#### 3. EXPERIMENTS AND DISCUSSIONS

#### 3.1. Evaluation Metrics

To quantitatively assess the performance of different algorithms, we introduce the evaluation metrics Recall

rate (R, 17), Precision rate (P, 18),  $F_1$  – score  $(F_1, 19)$ , location error  $(\Delta X, 20)$ , flux fluctuation  $(\Delta F lux, 21)$ , and regional intersection-over-union (IOU, 22) (Jiang et al. 2022).

$$R = \frac{TP}{TP + FN} \tag{17}$$

$$P = \frac{TP}{TP + FP} \tag{18}$$

$$F_1 = \frac{2TP}{2TP + FP + FN} \tag{19}$$

$$\Delta X = \frac{1}{N} \sum_{i}^{N} ||Cen_{d_{i}} - Cen_{s_{i}}||, i = 1, 2, ..., N \quad (20)$$

$$\Delta Flux = \frac{1}{N} \sum_{i}^{N} \frac{Sum_{d_{i}} - Sum_{s_{i}}}{Sum_{s_{i}}}, i = 1, 2, ..., N \quad (21)$$

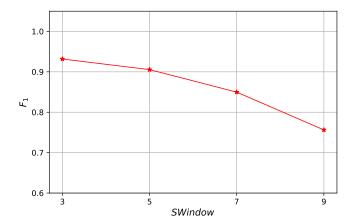
$$IOU = \frac{1}{N} \sum_{i}^{N} \frac{Region_{d_{i}} \cap Region_{s_{i}}}{Region_{d_{i}} \cup Region_{s_{i}}}, i = 1, 2, ..., N$$
(22)

where TP is the number of clumps detected correctly, FN is the number of missed clumps, and FP is the number of clumps detected wrongly. Cen represents the clump center, Sum represents the flux of a clump, Region represents the region of a clump, and N is the number of clumps detected correctly.  $s_i$  is the ith simulated parameter, and  $d_i$  refers to the corresponding parameter that is detected correctly. If  $\Delta X$  between a detected clump and a simulated clump is not greater than 2 voxels, the clump is considered to be correctly detected.

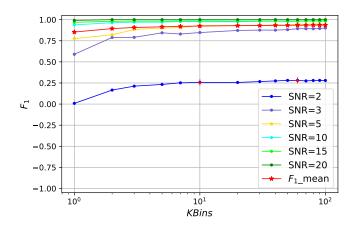
The metrics R, P, and  $F_1$  are utilized to assess the completeness and precision of the detections.  $\Delta X$  and  $\Delta Flux$  reflect the location error and flux loss of the sources, respectively. IOU measures the similarity between the simulated regions and the detected regions, which is crucial for calculating physical parameters, analyzing the morphological characteristics, and other applications. The Signal-to-Noise Ratio (SNR) is defined as the ratio of peak to the noise level, and the error bar is the standard deviation of the statistic in each SNR interval.

# 3.2. Experiments with simulated clumps 3.2.1. Simulated clumps

The model presented in Jiang et al. (2022) is adopted to generate 100 PPV data cubes with  $100 \times 100 \times 100$  voxels and each data cube contains 100 Gaussian clumps.



**Figure 11.** The variation trend graph of the mean value of  $F_1$  with respect to SWindow.



**Figure 12.** The variation trend graph of  $F_1$  with respect to KBins under different SNRs. The red curve is the mean value of  $F_1$  for all SNRs.

Gaussian noise with a RMS of 0.22 K is added to mimic realistic observational data. The peak intensities range from 0.44 to 4.4 K, ensuring a distribution of SNR between 2 and 20. Sizes vary between 2 and 4 voxels and rotation angles vary randomly. The distance between the truncated edge and the clump center is 3 times the sizes of the clump.

# 3.2.2. Determine appropriate values for the parameters of FacetClumps

To determine appropriate parameter values for FacetClumps, we investigate the influence of two parameters, namely SWindow (the scale of the window function) and KBins (the coefficient used to calculate the number of eigenvalue bins), on  $F_1$  in the simulated clumps. In addition, we also consider the instrument-related parameter SRecursionLBV, which is discussed in Appendix B. The simulated clumps generated by the

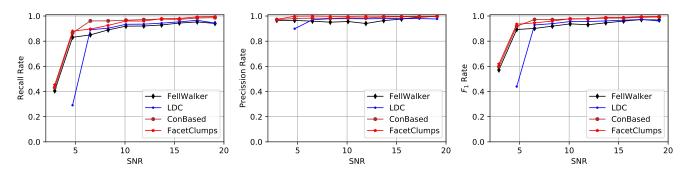


Figure 13. The statistics of R (left), P (middle), and  $F_1$  (right) curvers of FellWalker, LDC, ConBased, and FacetClumps for the simulated clumps. Black being FellWalker, blue being LDC, brown being ConBased and red being FacetClumps.

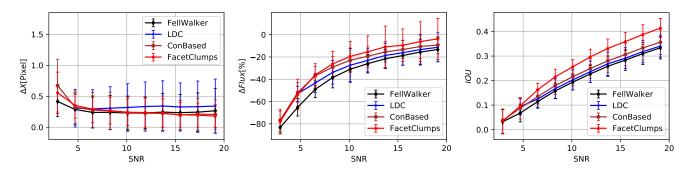


Figure 14. The statistics of  $\Delta X$  (left),  $\Delta Flux$  (middle), and IOU (right) curvers of FellWalker, LDC, ConBased, and FacetClumps for the simulated clumps.

simulation method proposed by Jiang et al. (2022) represent the theoretically detectable limit of all algorithms mentioned in this paper, i.e., these algorithms are only possible to detect clumps with valid peaks. Thus these experiments serve as an effective reference for choosing parameter values. As  $F_1$  is a comprehensive metric that evaluates both recall and precision rates, it is used as the evaluation metric in this study.

Figure 11 shows the variation trend of the average  $F_1$  with respect to the parameter SWindow. As SWindow increases from 3 to 9,  $F_1$  gradually decreases from 0.932 to 0.756, indicating that a smaller SWindow value leads to better detection performance. Therefore, the default value of SWindow is 3 (the minimum scale), and the recommended values are 3, 5, and 7. When dealing with poor quality observational data or large target sources, it is advisable to increase the value of SWindow.

Figure 12 shows the variation trend of  $F_1$  with respect to the parameter KBins under different SNRs. KBins varies from 1 to 100. The red line is the mean value of  $F_1$ , which ranges from 0.853 to 0.936. When SNR is less than 5,  $F_1$  is much lower than the average value, and the curve is a bit volatile, indicating that KBins has a small effect on the detection of faint clumps. When KBins is greater than 10 and SNR is greater than 5,  $F_1$  are almost straight lines, indicating that the change of KBins has

a little effect on the detection of clumps with high SNR. The mean value of  $F_1$  increases gradually with the increase of KBins, and then tends to plateau, indicating that larger KBins detect more clumps correctly.

As KBins increases from 1 to 10, there is a significant increase in  $F_1$ , which then gradually stabilizes as KBins continues to increase. As KBins increases,  $F_1$  for SNR of 2 first increases and then decreases, with the turning point occurring when KBins reaches 60. Therefore, the recommend values of KBins range from 10 to 60, with a default value of 35. The recommended and default values for all parameters are shown in appendix Table 9.

#### 3.2.3. Compare with other algorithms in simulated clumps

The comprehensive performance of FellWalker, LDC, and ConBased are better than GaussClumps, ClumpFind and ReinHold (Watson 2010; Berry 2015; Li et al. 2020; Luo et al. 2022; Jiang et al. 2022), making them suitable as comparators for our study. The changes in each evaluation metric for different algorithms with respect to SNR are statistically analyzed and presented in Figure 13 and Figure 14.  $\Delta X$  and  $\Delta Flux$  of LDC are based on the direct detections, and the Multiple Gaussian Model (Luo et al. 2022) is not used in this paper.

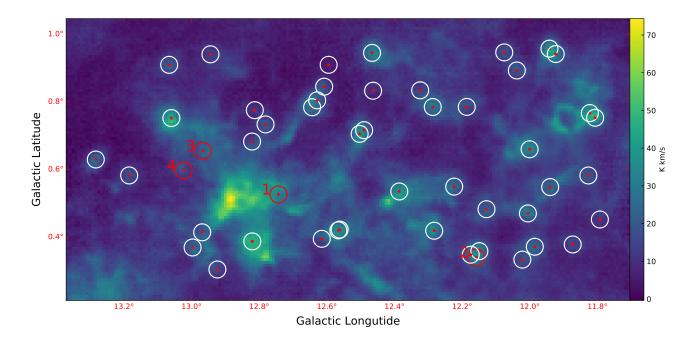


Figure 15. An example of the synthetic data of Data1. It is superimposed by the  $^{13}CO$  emission within  $11.7^{\circ} \le l \le 13.4^{\circ}$ ,  $0.22^{\circ} \le b \le 1.05^{\circ}$  and 5 km s<sup>-1</sup>  $\le v \le 35$  km s<sup>-1</sup> and simulated clumps. The cube size is  $180 \times 100 \times 200$  voxels. The number of simulated clumps is 50, evenly distributed in peak intensities ranging from 0.44 to 16.8 K, sizes ranging from 2 to 4 voxels, and with random angles. The red asterisks denote the central locations of the simulated clumps. The white circles denote the simulated clumps detected by FacetClumps and the red circles denote the missed simulated clumps. In total, 680 clumps are detected.

The left, middle, and right panels of Figure 13 show R, P, and  $F_1$ , respectively. R and  $F_1$  for each algorithm increase gradually with the increase of SNR. R of FacetClumps ranges from 0.453 to 0.996, with a mean value of 0.902, which is greater than that of FellWalker and LDC. P of FacetClumps varies between 0.973 to 1, with a mean value of 0.996.  $F_1$  of FacetClumps ranges from 0.619 to 0.997, with a mean value of 0.939. P and  $F_1$  of FacetClumps are greater than those of FellWalker, LDC and ConBased, and have minimal fluctuations.

The left, middle, and right panels of Figure 14 show  $\Delta X$ ,  $\Delta Flux$ , and IOU, respectively. It can be seen that clumps with lower SNR have higher error in locations, larger loss in measured fluxes, and lower IOU.  $\Delta X$  of FacetClumps decreases from 0.56 to 0.18 voxel. The average  $\Delta X$  of FacetClumps is similar to that of FellWalker, and less than that of ConBesed and LDC.  $\Delta Flux$  of FacetClumps increases from -78% to -3.7%, with a flux loss less than other algorithms. The low bias of the fluxes can be corrected using methods such as extrapolation proposed by Rosolowsky & Leroy (2006). IOU of FacetClumps increases from 0.03 to 0.41, which is higher than that of other algorithms.

In summary, FacetClumps exhibits greater  $F_1$ , indicating its better anti-noise performance; its flux loss is smaller and IOU is higher, demonstrating its superior capacity for both the flux recovery of clumps and the segmentation of overlapping clumps.

#### 3.3. Experiments with observational data

#### 3.3.1. The observational and synthetic data

The Milky Way Imaging Scroll Painting (MWISP) survey (Su et al. 2019) conducted by the Purple Mountain Observatory (PMO) is a large-field survey of  $^{12}CO$ ,  $^{13}CO$ , and  $C^{18}O$  (J=1-0) emission lines. The optical thickness of  $^{13}CO$  (J=1-0) is generally lower than that of  $^{12}CO$  (J=1-0), and the chemical properties of the isotope molecule itself are stable.  $C^{18}O$  haves a lower abundance than  $^{13}CO$ , and the optical thickness of  $C^{18}O$  (J=1-0) in the same line of sight is lower than that of the  $^{13}CO$  (J=1-0) transition lines, making it an ideal tool for detecting regions of higher density.

The half power beam width (HPBW) of  $^{13}CO$  and  $C^{18}O$  emission in MWISP is 52", with a grid spacing of 30". However, due to undersampling, the spatial sampling rate is set to FwhmBeam = 2. Similarly, the velocity sampling rate is VeloRes = 2. The spectral

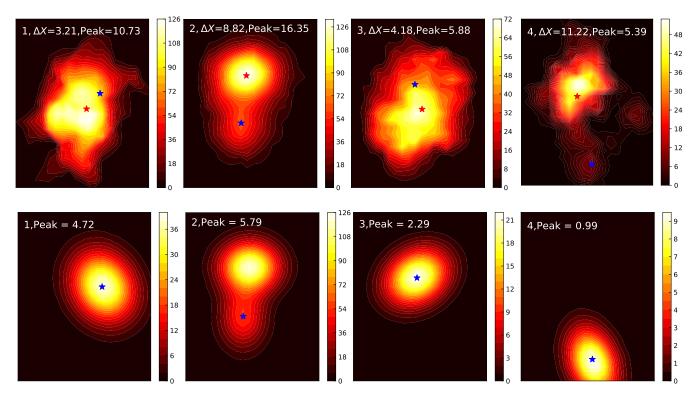


Figure 16. The upper panels show the nearest detected clumps of the missed clump centers shown in Figure 15. The lower panels show the undetected simulated clumps. The blue asterisks denote the location of the simulated clumps and the red asterisks denote the location of the closest detected clumps.  $\Delta X$  is the distance between the two locations in the graph. The unit of intensity is K.

**Table 3.** The average number of clumps, the SNR-weighted average R of the varying SNR, the corresponding flux and SNR when the R is equal to 0.9, the corresponding flux and SNR when R is equal to 0.8, the SNR-weighted average  $\Delta X$  in the spatial direction and in the velocity channels, the SNR-weighted average  $\Delta F lux$  and IOU.

Algorithm	N	$R_{mean}(SNR)$	$R_{0.9}(Flux/SNR)$	$R_{0.8}(Flux/SNR)$	$\Delta X_{LB}$	$\Delta X_V$	$\Delta F lux$	IOU
FellWalker	546	74.4%	-/-	560/30	0.31	0.2	10.7%	0.42
LDC	692	80.3%	-/-	180/15	0.33	0.19	9.4%	0.41
ConBased	657	83.1%	400/30	170/10	0.27	0.18	10.7%	0.44
FacetClumps	671	90.2%	190/14	100/8	0.17	0.12	30.7%	0.5

resolution of the observational data is approximately 0.166 km s<sup>-1</sup>, and the total fluxes are multiplied by this factor to obtain physically meaningful values. The  $^{13}CO$  and  $C^{18}O$  (J=1-0) emission of MWISP within  $11.7^{\circ} \leq l \leq 13.4^{\circ},~0.22^{\circ} \leq b \leq 1.05^{\circ}$  and 5 km s<sup>-1</sup>  $\leq v \leq 35$  km s<sup>-1</sup> where many notable star-forming activities have been discovered (e.g. Li et al. 2013; Chen et al. 2016, 2017) are used to examine the performance of FacetClumps. The noise level of the  $^{13}CO$  and  $C^{18}O$  emissions are approximately 0.22 K and 0.2 K, respectively. The maximum intensity of the  $^{13}CO$  emission is approximately 16.8 K.

We generate 200 synthetic data cubes named Data1, which are superimposed with the  $^{13}CO$  emission and

simulated clumps. Each synthetic data cube contains 50 simulated clumps, with uniform peak intensities ranging from 0.44 ( $2 \times RMS$ ) to 16.8 K (maximum intensity) to ensure that the statistical magnitude of clumps is consistent under different SNRs, uniform sizes ranging from 2 to 4 voxels, and random angles. The flux is primarily distributed between 20 and 1200 K km s<sup>-1</sup>, which closely resembles the distribution observed in actual molecular clumps.

To evaluate the performance of the FacetClumps in a wider range of signal densities and environments, we have selected a high-density and a low-density area of  $^{13}CO$  emission to create more diverse synthetic datasets (Data2 and Data3, see Appendix A for the details). In

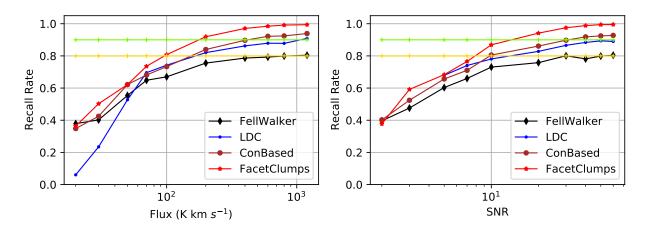
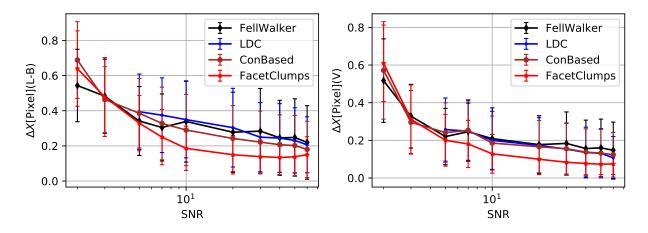


Figure 17. The statistics of R of Fellwalker, LDC, ConBased, and FacetClumps for the synthetic data. The left panel shows R as a function of flux, and the right panel shows R as a function of SNR. The lawngreen line is equal to 0.9, and the gold line is equal to 0.8.



**Figure 18.** The statistics of  $\Delta X$  of FellWalker, LDC, ConBased, and FacetClumps for the synthetic data.  $\Delta X$  is a function of SNR. The left panel shows  $\Delta X$  in the spatial direction, and the right panel shows  $\Delta X$  in the velocity channels.

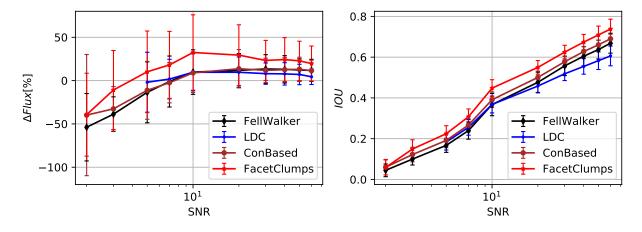


Figure 19. The statistics of  $\Delta Flux$  (left) and IOU (right) of FellWalker, LDC, ConBased and FacetClumps for the synthetic data.  $\Delta Flux$  and IOU are functions of SNR.

addition, to investigate the robustness of FacetClumps parameters under various sampling conditions, we have resampled the synthetic data of Data1 using different sampling factors in different directions (see Appendix B for the details).

An example of the synthetic data of Data1 is shown in Figure 15. The red asterisks in the white circle denote simulated clumps detected by FacetClumps, while the red circles denote missed simulated clumps. Figure 16 shows the four missed clumps, where the blue asterisks denote the locations of the simulated clumps and the red asterisks denote the locations of the closest detected clumps. It can be seen that the undetected simulated clumps usually have large overlaps with real clumps of higher intensity (e.g., Nos. 1 and 3), are affected by other simulated clumps (e.g., Nos. 2), or have a low SNR and have become part of real ones (e.g., Nos. 4).

#### 3.3.2. Compare with other algorithms in synthetic data

To test the performance of the algorithms in observational environments, we apply them to detect the synthetic data Data1. The average number of detected clumps by FellWalker, LDC, ConBased, and FacetClumps is 546, 692, 657, and 671, respectively. We analyze several evaluation metrics, including R as a function of flux and SNR,  $\Delta X$  in the spatial direction and in the velocity channels as a function of SNR, and  $\Delta Flux$  and IOU as functions of SNR. The results of these analyses are shown in Figure 17, Figure 18, and Figure 19, respectively.

When computing the average metrics, we take into account the weighted mean value of the intervals with SNR greater than 5. The weights are obtained by using FellWalker to detect the observational data used to construct the synthetic data, and then calculating the corresponding weights of the SNR distribution for each interval. The average number of clumps, the SNR-weighted average R of the varying SNR, the corresponding flux and SNR when R is equal to 0.9 and 0.8, the SNR-weighted average  $\Delta X$  in the spatial direction and in the velocity channels, and the SNR-weighted average  $\Delta Flux$  and IOU are presented in Table 3.

As shown in Figure 17, R of FacetClumps ranges from 0.38 to 0.99, which is higher than that of other algorithms, especially when flux and SNR are low. This indicates that FacetClumps detects more correct clumps and can better detect faint clumps in observational background. When R reaches 0.9, the corresponding flux/SNR of ConBased and FacetClumps are about 400 K km s<sup>-1</sup>/30 and 190 K km s<sup>-1</sup>/14, respectively. When R reaches 0.8, the corresponding flux/SNR of FellWalker, LDC, ConBased and FacetClumps are

**Table 4.** The number of clumps, the cross-matching rate of  $C^{18}O/^{13}CO$ , and the relative time unit of the  $^{13}CO$  emission.

Algorithm	$^{13}CO$	$C^{18}O$	$C^{18}O/^{13}CO$	$T_{^{13}CO}$
FellWalker	538	134	48.5%	/
LDC	680	138	62.3%	3.2
ConBased	629	105	68.6%	1.9
FacetClumps	692	185	56.8%	1

about 560 K km s<sup>-1</sup>/30, 180 K km s<sup>-1</sup>/15, 170 K km s<sup>-1</sup>/10 and 100 K km s<sup>-1</sup>/8, respectively. R of Con-Based is higher than that of FellWalker and LDC, and R of LDC is higher than that of FellWalker. The SNR-weighted average R of FellWalker, LDC, ConBased, and FacetClumps is 74.4%, 80.3%, 83.1%, and 90.2%, respectively.

The left panel of Figure 18 shows  $\Delta X$  in the spatial direction, while the right panel shows  $\Delta X$  in the velocity channels.  $\Delta X$  of FacetClumps in the spatial direction decreases from 0.64 to 0.14 voxel, with a SNR-weighted mean value of 0.17 voxel, and that in the velocity channels decreases from 0.61 to 0.07 voxel, with a SNR-weighted mean value of 0.12 voxel.  $\Delta X$  of FacetClumps in the spatial direction and in the velocity channels are smaller, indicating that the clumps detected by FacetClumps in the real environment have more precise locations.

 $\Delta Flux$  and IOU are shown in the left and right panels of Figure 19, respectively.  $\Delta Flux$  of FacetClumps is between -39.4% and 32.4%. The SNR-weighted average  $\Delta Flux$  of FellWalker, LDC, and ConBased is 10.7%, 9.4%, 10.7%, and 30.7%, indicating that FacetClumps has less flux loss. IOU of FacetClumps increases from 0.06 to 0.74, and its SNR-weighted mean value is 0.5. The SNR-weighted average IOU of FellWalker, LDC, and ConBased are 0.42, 0.41, and 0.44, indicating that FacetClumps can better segment the regions of simulated clumps from real signals.

In summary, FacetClumps exhibits a greater R and a smaller  $\Delta X$ , indicating that it is better suited for the observational environments and can locate clumps more accurately in complex backgrounds. The  $\Delta Flux$  and IOU statistics show that FacetClumps performs slightly better in detecting useful signals and segmenting different clumps. Furthermore, the R curve of FacetClumps is smoother and the error bars for  $\Delta X$  are shorter, indicating its superior stability.

#### 3.3.3. Experiments to observational data

To evaluate the usability of FacatClumps on observational data, we apply FellWalker, LDC, ConBased, and FacetClumps to detect the data cubes of  $^{13}CO$ 

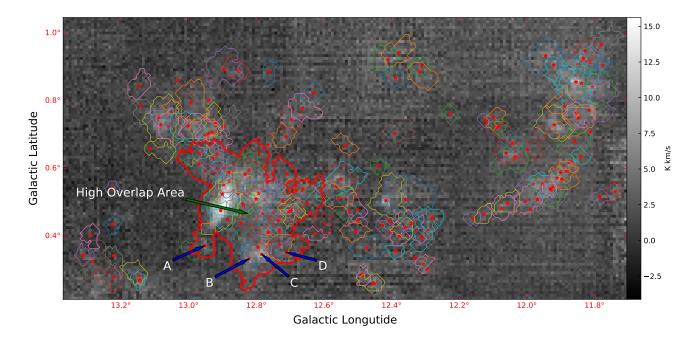
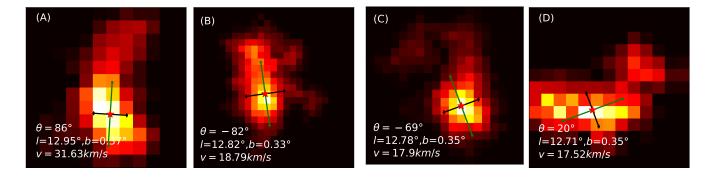


Figure 20. The result of the application of FacetClumps in observational data of  $C^{18}O$  emission within  $11.7^{\circ} \le l \le 13.4^{\circ}$ ,  $0.22^{\circ} \le b \le 1.05^{\circ}$  and 5 km s<sup>-1</sup>  $\le v \le 35$  km s<sup>-1</sup>. The total number of clumps is 185, of which 163 did not touch the edge. The red asterisks denote the central locations of the clumps which do not touch the edge, and different thin outlines delineate the boundaries of different clumps. The thick red outline circles an area of high overlap, with four clumps (A, B, C, and D) at the bottom of the area, as shown in Figure 21, and the other clumps are shown in Figures 27 and 28.



**Figure 21.** Velocity-integrated intensity images of four clumps in the high overlap area as shown in Figure 20. A, B, C, and D correspond to the markers in Figure 20. The red asterisks denote the central locations of the clumps. The green lines denote the principal axis and the black lines denote the secondary axis. The ratio of the lengths of the principal and secondary axes is equivalent to the ratio of their respective axes.  $\theta$  denotes the angle between the principal axis and the negative direction of galactic longitude. (l, b, v) is the central coordinate.

and  $C^{18}O$  emission. The number of clumps, the crossmatching rate, and the relative time unit T in  $^{13}CO$ emission are presented in Table 4. The cross-matching rate of Line1/Line2 is defined as the percentage of Line2 clumps that coincide with Line1 clumps (Li et al. 2020; Jiang et al. 2022). The results show that the crossmatching rate of FacetClumps is higher than that of Fell-Walker, and ConBased has the highest cross-matching rate. Since the time spent in different programming languages is not comparable, the minimum time spent by the algorithm in the same language is recorded as 1 unit, and the time spent by the other algorithms in the same language is recorded as a multiple of this unit. LDC, ConBased, and FacetClumps use the same programming language, while FellWalker uses a different one. As shown in Table 4, the time taken by LDC and ConBased is 3.2 and 1.9 times that of FacetClumps, respectively.

Figure 20 shows the results of the application of FacetClumps in the  $C^{18}O$  emission. A total of 185 clumps are identified, of which 163 do not touch the edges. The red asterisks denote the central locations of the clumps that do not touch the edges. Different thin outlines delineate the boundaries of different clumps, while the thick red outline circles an area of high overlap. A, B, C, and D represent the four clumps at the bottom of the high overlap area, as shown in Figure 21, and the other clumps are shown in Figures 27 and 28. Most clumps contain a well-defined denser central source surrounded by weaker gas, while some may be dual-clump systems (e.g., Figure 21(B)), or some may have trailing substructures (e.g., Figure 21(C)). These images illustrate that FacetClumps effectively and accurately detect clumps, even in areas with a relatively higher degree of overlap. Further scientific analysis of the clumps will be carried out in the future.

#### 4. SUMMARY

We propose FacetClumps, a new algorithm for detecting molecular clumps in astronomical data. Initially, the signal regions are extracted based on morphology. Then, the Gaussian Facet model operators are utilized to fit the signal regions to derive the first and second derivatives, which are used in conjunction with the maximum determination theorem to locate the clump centers. Subsequently, signal regions are segmented into local regions

based on local gradients. Finally, the local regions are clustered into the clump centers based on connectivity and minimum distance, thus identifying the region of each clump.

We have conducted parametric and comparative experiments on simulated clumps with different SNRs to determine appropriate values for the instrument-independent parameters of FacetClumps and to evaluate the performance of Fellwalker, LDC, ConBased, and FacetClumps. The parametric experiments provide an effective reference for the parameters, while the comparative experiments demonstrate that FacetClumps improves the noise resistance and segmentation accuracy of overlapping clumps even further.

We have conducted a series of experiments with synthetic data. Experiments performed in synthetic data consisting of an active star-forming zone and simulated clumps demonstrate that FacetClumps has greater R with 90.2%, smaller  $\Delta X$ , better  $\Delta F lux$ , and higher IOU. Experiments performed in two types of larger synthetic data with different signal densities in Appendix A demonstrate that FacetClumps has pronounced advantages in high-density signal environments. Experiments performed in different resampled synthetic data in Appendix B indicate that the parameters of FacetClumps have self-adaptability. We have conducted some tests in the observational data, which illustrate that FacetClumps is more efficient and can be applied in observational data well.

## ACKNOWLEDGEMENTS

We are grateful to the anonymous referees for their invaluable insights and comments, which enabled us to refine and enhance this work. This work is supported by the National Natural Science Foundation of China (grants Nos. U2031202, 11903083, 11873093). This research make use of the data from the Milky Way Imaging Scroll Painting (MWISP) project, which is a multi-line survey in  ${}^{12}CO/{}^{13}CO/{}^{C18}O$  along the northern galactic plane with PMO-13.7m telescope. We are grateful to all the members of the MWISP working group, particularly the staff members at PMO-13.7m telescope, for their long-term support. MWISP is sponsored by National Key R&D Program of China with grant 2017YFA0402701 and CAS Key Research Program of Frontier Sciences with grant QYZDJ-SSW-SLH047.

#### REFERENCES

Alves, J., Lombardi, M., & Lada, C. J. 2007, A&A, 462,

 $L17,\ doi:\ 10.1051/0004\text{-}6361\text{:}20066389$ 

An, X. 2007, Journal of Computer-Aided Design & Computer Graphics

Berry, D. S. 2013, Starlink User Note, 255

- —. 2015, Astronomy and Computing, 10, 22, doi: 10.1016/j.ascom.2014.11.004
- Blitz, L., & Shu, F. H. 1980, ApJ, 238, 148, doi: 10.1086/157968
- Blitz, L., & Williams, J. P. 1999, arXiv e-prints, astro. https://arxiv.org/abs/astro-ph/9903382
- Boomgaard, V., & Weijer, V. 2003, in International Conference on Scale-Space Theories in Computer Vision
- Brejl, M., & Sonka, M. 2000, Comput. Vis. Image Underst., 77, 84
- Chabrier, G. 2003, PASP, 115, 763, doi: 10.1086/376392
- Chen, S., & Haralick, R. 1995, IEEE Transactions on Image Processing, 4, 335, doi: 10.1109/83.366481
- Chen, Z., Jiang, Z., Tamura, M., Kwon, J., & Roman-Lopes, A. 2017, ApJ, 838, 80, doi: 10.3847/1538-4357/aa65d3
- Chen, Z., Zhang, S., Zhang, M., et al. 2016, ApJ, 822, 114, doi: 10.3847/0004-637X/822/2/114
- Dame, T. M., Hartmann, D., & Thaddeus, P. 2001, ApJ, 547, 792, doi: 10.1086/318388
- Dame, T. M., Ungerechts, H., Cohen, R. S., et al. 1987, ApJ, 322, 706, doi: 10.1086/165766
- Gammie, C. F., Lin, Y.-T., Stone, J. M., & Ostriker, E. C. 2003, ApJ, 592, 203, doi: 10.1086/375635
- Haralick, R. M. 1984, IEEE Transactions on Pattern Analysis & Machine Intelligence, PAMI-6, 58
- Jiang, Y., Zheng, S., Jiang, Z., et al. 2022, Astronomy and Computing, 100613,
  - doi: https://doi.org/10.1016/j.ascom.2022.100613
- Jiang,, Y., 2023, FacetClumps: A Facet-based Molecular Clump Detection Algorithm, Version 0.0.4, Zenodo, doi: 10.5281/zenodo.7991006
- Koda, J., Sawada, T., Hasegawa, T., & Scoville, N. Z. 2006, ApJ, 638, 191, doi: 10.1086/498640
- Koskinen, L., Astola, J. T., & Neuvo, Y. A. 1991, in
  Society of Photo-Optical Instrumentation Engineers
  (SPIE) Conference Series, Vol. 1568, Image Algebra and
  Morphological Image Processing II, ed. P. D. Gader &
  E. R. Dougherty, 262–270, doi: 10.1117/12.46121
- Lee, J.-W., Carney, B. W., Fullton, L. K., & Stetson, P. B. 2001, AJ, 122, 3136, doi: 10.1086/324461
- Li, C., Wang, H., Zhang, M., et al. 2018, ApJS, 238, 10, doi: 10.3847/1538-4365/aad963
- Li, C., Wang, H.-C., Wu, Y.-W., Ma, Y.-H., & Lin, L.-H. 2020, Research in Astronomy and Astrophysics, 20, 031, doi: 10.1088/1674-4527/20/3/31

- Li, J.-Y., Jiang, Z.-B., Liu, Y., & Wang, Y. 2013, Research in Astronomy and Astrophysics, 13, 921, doi: 10.1088/1674-4527/13/8/004
- Lodieu, N. 2013, MNRAS, 431, 3222, doi: 10.1093/mnras/stt402
- Luo, X., Zheng, S., Huang, Y., et al. 2022, Research in Astronomy and Astrophysics, 22, 015003, doi: 10.1088/1674-4527/ac321d
- Narayanan, D., Cox, T. J., Shirley, Y., et al. 2008, ApJ, 684, 996, doi: 10.1086/588720
- Narayanan, D., Cox, T. J., Shirley, Y. L., et al. 2007, ApJ, 684, 996, doi: 10.1086/321609
- Qiang, J., & Haralick, R. M. 2002, Pattern Recognition, 35, 689
- Rosolowsky, E., & Leroy, A. 2006, PASP, 118, 590, doi: 10.1086/502982
- Sanders, D. B., Clemens, D. P., Scoville, N. Z., & Solomon, P. M. 1986, ApJS, 60, 1, doi: 10.1086/191086
- Serra, J. 1982, Image analysis and mathematical morphology (Academic Press, New York)
- Shu, F. H.-S., Adams, F. C., & Lizano, S. 1987, Annual Review of Astronomy and Astrophysics, 25, 23
- Sinha, D., & Dougherty, E. R. 1992, in Society of Photo-Optical Instrumentation Engineers (SPIE)
  Conference Series, Vol. 1607, Intelligent Robots and Computer Vision X: Algorithms and Techniques, ed.
  D. P. Casasent, 477–488, doi: 10.1117/12.57084
- Stutzki, J., & Guesten, R. 1990, ApJ, 356, 513, doi: 10.1086/168859
- Su, Y., Yang, J., Zhang, S., et al. 2019, ApJS, 240, 9, doi: 10.3847/1538-4365/aaf1c8
- Testi, L., & Sargent, A. I. 1998, ApJL, 508, L91, doi: 10.1086/311724
- Watson, M. 2010, Master's thesis, University of Hertfordshire, UK
- Williams, J. P., de Geus, E. J., & Blitz, L. 1994, ApJ, 428, 693, doi: 10.1086/174279
- Wu, K., Otoo, E., & Shoshani, A. 2005, in Society of Photo-Optical Instrumentation Engineers (SPIE)
  Conference Series, Vol. 5747, Medical Imaging 2005: Image Processing, ed. J. M. Fitzpatrick & J. M. Reinhardt, 1965–1976, doi: 10.1117/12.596105
- Wu, K., Otoo, E. J., & Suzuki, K. 2008, Pattern Analysis and Applications, 12, 117
- Wu, X., & Wee, W. G. 1999, Spie Proceedings, 3586, 319

#### APPENDIX

#### A. COMPARATION OF DIFFERENT ALGORITHMS ON A LARGER RANGE OF SYNTHETIC DATA

We select the  $^{13}CO$  (J=1-0) emission of MWISP within  $13^{\circ} \le l \le 16^{\circ}$ ,  $-1.5^{\circ} \le b \le 0.5^{\circ}$  and  $0 \text{ km s}^{-1} \le v \le 70 \text{ km s}^{-1}$  and within  $184.5^{\circ} \le l \le 187.5^{\circ}$ ,  $-1^{\circ} \le b \le 1^{\circ}$  and  $-10 \text{ km s}^{-1} \le v \le 60 \text{ km s}^{-1}$  as high-density and low-density signal environments, respectively. They are used to construct synthetic data, namely Data2 and Data3, as shown in Figure 22. Each dataset contains 100 synthetic data cubes, with each cube containing 100 simulated clumps. The range of peak intensities is from  $2 \times RMS$  to maximum intensity, and the range of sizes is from 1 to 5 voxels.

To evaluate the performance of completeness and the accuracy of locations for the algorithms under different densities, we analyse the variations of R with flux and SNR and  $\Delta X$  in PPV with SNR, which are shown in Figure 23 and Figure 24, respectively. The average numbers of clumps, the SNR-weighted average R, and the SNR-weighted average  $\Delta X$  of the varying SNR are summarized in Table 5.

In the high-density signal environment: The average number of FacetClumps is 3769. FacetClumps has advantages in terms of R (92%) and  $\Delta X$  (0.36 voxel) compared to ConBased, LDC, and FellWalker; R (85.7%) and  $\Delta X$  (0.41 voxel) of ConBased is the second best only to FacetClumps; R (84.9%) of LDC is better than R (79.9%) of FellWalker, and  $\Delta X$  (0.46 voxel) of LDC is greater than  $\Delta X$  (0.45 voxel) of FellWalker. In the low-density signal environment: The average number of FacetClumps is 290. FacetClumps has an advantage over FellWalker and LDC in terms of R (98.8%); R (98.8%) of ConBased is higher than that of FellWalker and LDC; R (93.6%) of LDC is better than R (89.6%) of FellWalker.  $\Delta X$  obtained by FacetClumps (0.28 voxel) and ConBased (0.27 voxel) is greater than that of FellWalker (0.26 voxel) and LDC (0.26 voxel).

In brief, FacetClumps has obvious advantages over other algorithms in high-density signal environments.

**Table 5.** The average number of clumps, the SNR-weighted average R, and the SNR-weighted average  $\Delta X$  of the varying SNR of Data2 and Data3.

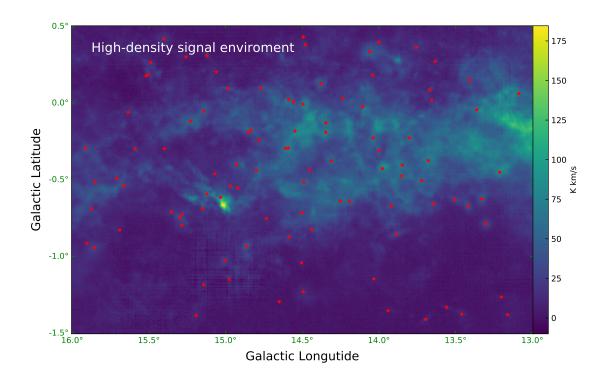
Algorithm	$N_{Data2}$	$N_{Data3}$	$R_{Data2}$	$R_{Data3}$	$\Delta X_{Data2}$	$\Delta X_{Data3}$
FellWalker	3557	215	79.9%	89.6%	0.45	0.26
LDC	4492	202	84.9%	93.6%	0.46	0.26
ConBased	4776	237	85.7%	98.8%	0.41	0.27
FacetClumps	3769	290	92%	98.8%	0.36	0.28

#### B. CONFIGURATION PARAMETERS OF FACETCLUMPS UNDER DIFFERENT RESAMPLE FACTORS

Downsampling increase the overlap among clumps, making the peaks of some clumps disappear and become undetectable. To roughly assess the effect of downsampling on the valid peaks, we apply FacetClumps to detect the  $^{13}CO$  emission within  $11.7^{\circ} \le l \le 13.4^{\circ}$ ,  $0.22^{\circ} \le b \le 1.05^{\circ}$  and 5 km s<sup>-1</sup>  $\le v \le 35$  km s<sup>-1</sup> to obtain the attributes of the clumps, which are used to generate a simulated data cube with the same volume and similar properties as the observed one. The simulated data is downsampled in spectral channels (hereafter V), spatial direction (hereafter V), and simultaneously in both spatial direction and spectral channels (hereafter V) with a sampling factor of 2, respectively. We check the change in the number of peaks in each resampled data. The results show that the number of detectable clumps decreased by about 3.5%, 6.7%, and 9.7%, respectively. It should be noted that these estimates only give a lower limit for reductions of detectable clumps in observed data, as some peaks in the original simulated data already have disappeared and downsampling makes it easier to eliminate peaks in irregular clumps.

To enable FacetClumps to be applied to multiple datasets, associating the parameters of FacetClumps with the parameters of the instrument will minimize systematic differences. In Section 3.3.2, we have discussed the influence of instrument-independent parameters SWindow and KBins on the performance of FacetClumps. Here, we discuss the associations between the instrument-related parameter SRecursionLBV and the FWHM of the instrument beam FwhmBeam and the velocity resolution of the instrument VeloRes, which are expressed by equation (B1). Besides, the associations between resampling and FwhmBeam and VeloRes are expressed by equation (B2).

$$SRecursionLB = (F_0 + FwhmBeam)^2, SRecursionV = V_0 + VeloRes$$
 (B1)



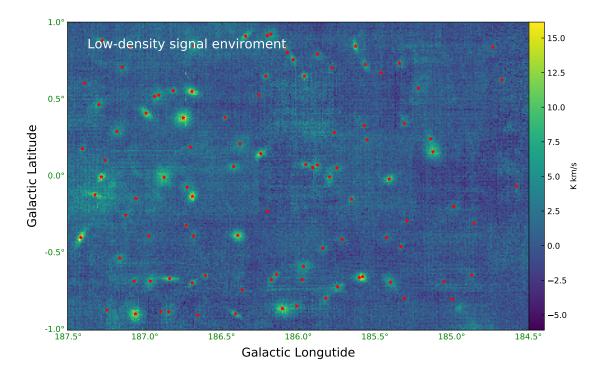


Figure 22. Examples of the synthetic data of Data2 and Data3. The top panel is superimposed by the  $^{13}CO$  emission within  $13^{\circ} \le l \le 16^{\circ}$ ,  $-1.5^{\circ} \le b \le 0.5^{\circ}$  and 0 km s<sup>-1</sup>  $\le v \le 70$  km s<sup>-1</sup> and simulated clumps, representing a typical high-density signal environment. The bottom panel is superimposed by the  $^{13}CO$  emission within  $184.5^{\circ} \le l \le 187.5^{\circ}$ ,  $-1^{\circ} \le b \le 1^{\circ}$  and -10 km s<sup>-1</sup>  $\le v \le 60$  km s<sup>-1</sup> and simulated clumps, representing a typical low-density signal environment. The cube size is  $423 \times 241 \times 361$  voxels, and there are 100 simulated clumps. The red asterisks denote the central locations of the simulated clumps.

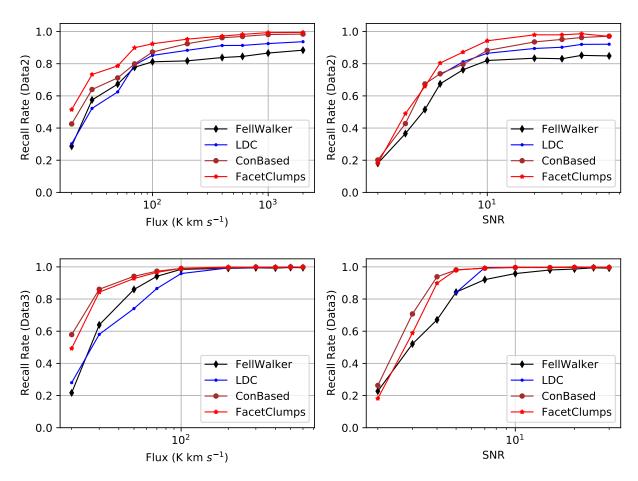


Figure 23. The statistics of R of FellWalker, LDC, ConBased, and FacetClumps for synthetic data of different density signal environments. The upper panels are the statistics of synthetic data in a high-density signal environment as shown in the top panel of Figure 22. The lower panels are the statistics of synthetic data in a low-density signal environment as shown in the bottom panel of Figure 22. The left panels depict R as a function of flux, and the right panels depict R as a function of SNR.

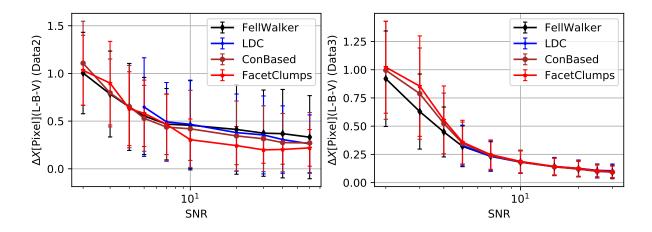


Figure 24. The statistics of  $\Delta X$  of FellWalker, LDC, ConBased, and FacetClumps for synthetic data of different density signal environments.  $\Delta X$  is a function of SNR. The left panel shows the statistic of synthetic data in a high-density signal environment as shown in the top panel of Figure 22. The right panel shows the statistic of synthetic data in a low-density signal environment as shown in the bottom panel of Figure 22.

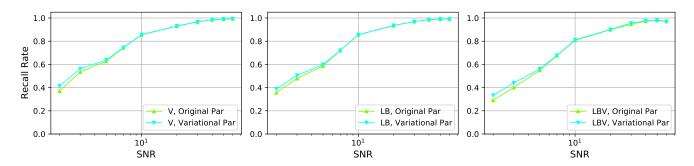


Figure 25. The statistics of R of resampling along different directions. V, LB, and LBV represent the statistics in data downsampled in the spectral channel, data downsampled in the spatial direction and data downsampled in both the spatial direction and the spectral channel, respectively. 'Original Par' means that SRecursionLBV uses the same value as when detecting unresampled synthetic data, and 'Variational Par' means that SRecursionLBV varies according to formula (B1).

$$FwhmBeam' = FwhmBeam \times SFactorLB, VeloRes' = VeloRes \times SFactorV$$
 (B2)

where, SFactorLB and SFactorV are the resample factors in spatial direction and spectral channels, respectively.  $F_0$  and  $V_0$  can be customized according to the minimum size of the clumps required for scientific objectives, with  $F_0 = 2$  and  $V_0 = 3$  as the default values. The values of FwhmBeam and VeloRes can be modified to suit the specific instrument parameters of the radio telescope used to acquire the molecular line data.

To explore the robustness of SRecursionLBV in different sampled data, we downsample the synthetic data of Data1 (FwhmBeam = 2, VeloRes = 2) described in Section 3.3.3 in V, LB, and LBV with a sampling factor of 2 (SFactorLB = 0.5, SFactorV = 0.5), respectively. The statistics of R of resampling along different directions are shown in Figure 25. SRecursionLBV adopts two different types of values, which are respectively the same value as when detecting unresampled synthetic data and the values varying with (B1). Figure 25 shows R of low SNR can be slightly improved by using variational parameters. Moreover, resampling can change the minimum size of clumps that are required, making the variational parameters even more desirable. Therefore, the following analysis is based on the statistics derived from the utilization of the variational parameters.

To further investigate the robustness of parameters, we have added a new similar downsampled dataset (SFactorLB=0.75, SFactorV=0.75), and the statistics of R of resampling with different resample factors and along different directions are shown in Figure 26. For SFactorLB=0.75, SFactorV=0.75: The reductions of the mean values of R are all less than 0.5%. For SFactorLB=0.5, SFactorV=0.5: When the SNR is greater than 10, the reductions of R in V sampling and LB sampling are less than 1%, and that in LBV sampling is less than 4%; The mean values of R in different resampled data decrease by about 0.8%, 1.2%, and 4% respectively, all of which are smaller than the corresponding reductions in the simulated data. Experimental results indicate that the parameters of FacetClumps can be effectively adapted to different sampled data, particularly for clumps with SNR of more than 10, where the parameters are reliable.

#### C. CONFIGURATION PARAMETERS OF DIFFERENT ALGORITHMS

The configuration parameters are presented in Table 6, Table 7, Table 8, and Table 9. For FacetClumps, we provide both default and recommended values.

#### D. PARTIAL CLUMPS IN THE HIGH OVERLAP AREA IN SECTION 3.3.3

Partial clumps in the high overlap area as described in Section 3.3.3 are shown in Figures 27 and 28. They are used to visually evaluate the usability of FacetClumps in detecting molecular clumps.

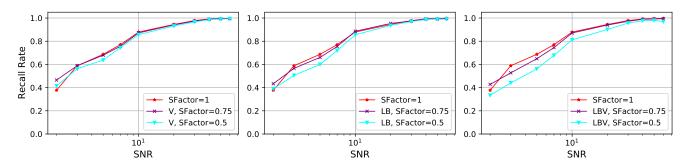


Figure 26. The statistics of R of resampling with different resample factors and along different directions. SFactor = 1 means that no resampling is applied to the data. V, SFactor = 0.5 means that the data is downsampled with a resample factor of 2 in the spectral channel, and other labels have similar meanings. The plots are the statistics of utilizing variational parameters.

#### Table 6. FellWalker parameters.

FellWalker.AllowEdge=1
FellWalker.CleanIter=1
FellWalker.FlatSlope=2\*RMS
FellWalker.FwhmBeam=2
FellWalker.MaxBad=0.05
FellWalker.MaxJump=4
FellWalker.MinDip=1\*RMS
FellWalker.MinHeight=3\*RMS
FellWalker.MinPix=27
FellWalker.Noise=2\*RMS
FellWalker.RMS=RMS
FellWalker.RMS=RMS
FellWalker.VeloRes=2

## Table 7. LDC parameters.

LDC.RMS=RMS LDC.Threshold=2\*RMS LDC.GradientMin=0.01 LDC.DistanceMin=4 LDC.PeakMin=5\*RMS LDC.PixelMin=27

## Table 8. ConBased parameters.

ConBased.RMS=RMS
ConBased.Threshold=2\*RMS
ConBased.RegionMin=27
ConBased.ClumpMin=216
ConBased.DIntensity=2\*RMS
ConBased.DDistance=8

 ${\bf Table~9.~FacetClumps~parameters.}$ 

 $FacetClumps.RMS {=} RMS$ 

FacetClumps.Threshold=2\*RMS,[n\*RMS]

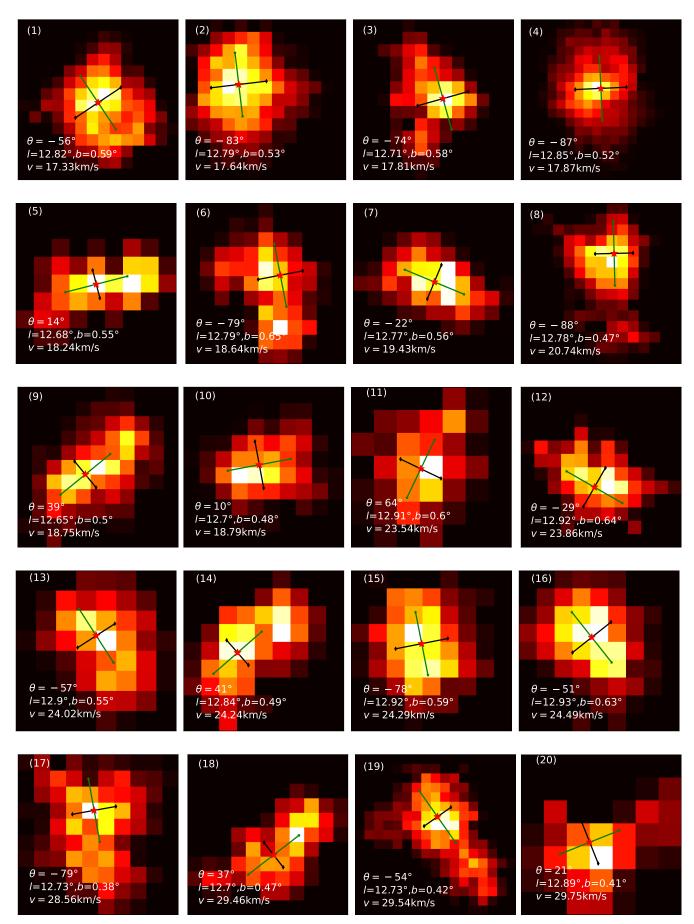
FacetClumps.SWindow=3,[3,5,7]

FacetClumps.KBins=35,[10,...,60]

 ${\it FacetClumps.FwhmBeam}{=}2$ 

 ${\it FacetClumps.VeloRes}{=}2$ 

 ${\it FacetClumps.} SRecursionLBV {=} [16,\!5]$ 



**Figure 27.** Velocity-integrated intensity images of clumps in the high overlap area as shown in Figure 20. Each panel is presented in the same manner as in Figure 21.

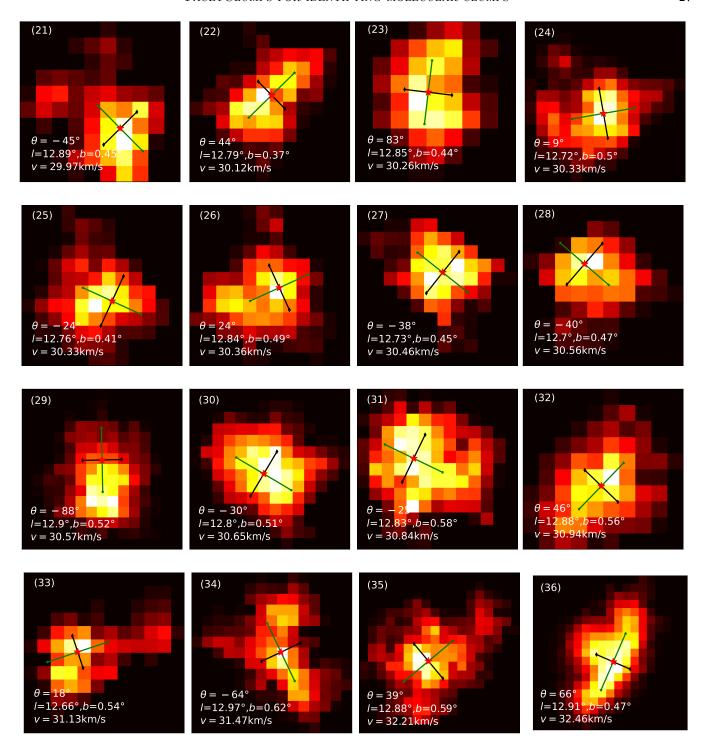


Figure 28. Velocity-integrated intensity images of clumps in the high overlap area as shown in Figure 20. Each panel is presented in the same manner as in Figure 21.